

A Realistic Unified Gauge Coupling from the

Micro-landscape of Orbifold GUTs

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Outline

- Two-Loop Casimir-Stabilization of $S^1 / (\mathbb{Z}_2 \times \mathbb{Z}'_2)$
[This a straight-forward extension of previous work with G.v. Gersdorff to a realistic SUSY model.]
- "Predicting" the 4d gauge coupling at M_{GUT}
- "Uplifting" the stabilized AdS_4 vacuum

Motivation

a) Field-Theoretic

- 5d SU_5 orbifold GUTs are the simplest realistic GUT models
- The 5d-Radius determines the 4d gauge coupling

b) String-Theoretic

- Heterotic orbifold models are among the most realistic string models of the "real world" (\rightarrow talks of M. Ratz, H. Nilles, J. Schmidt, ...)
- 5d orbifold GUTs can arise naturally as eff. theories between string- and GUT-scale
- Such "anisotropic orbifolds" may be helpful to overcome the string-scale / GUT-scale problem (Ibanez / Lüst '92 ; Witten '96
A.H. / Trappetti '04 ; Dundee / Raby / Wingerter '08)

The Problem

(We will take a purely 5d-field-theoretic point of view throughout)

$$\text{Compactification} \Rightarrow g_4^2 = \frac{g^2}{2\pi R}$$

Proper expansion
parameters:

$$\frac{g_4^2 N}{16\pi^2}$$

$$\frac{g^2 N}{24\pi^3} \equiv \frac{1}{M}$$

(fundamental scale)

The compactification - relation
takes the form:

$$\frac{g_4^2 N}{16\pi^2} = \frac{3}{4} \cdot \frac{1}{MR}$$

$\alpha_{\text{GUT}} \sim \frac{1}{25}$ implies $MR \approx 45$ ("mild hierarchy")

2-loop Casimir Stabilization

(cf. Gersdorff / A.H. '05)

$$S^1: \quad V(R) \sim \frac{1}{R^4} + \frac{g^2}{R^5} \sim \frac{1}{R^4} + \frac{1}{MR^5}$$

- The signs and values of coefficients follow from field content
- Ratio of coefficients determines numerical prefactor in

$$R_{\min} \sim \frac{1}{M}$$

- For appropriate field content, the prefactor can be (accidentally) large:

$$R_{\min} \gg \frac{1}{M}$$

- If so, higher loops are irrelevant and the minimum is "perturbatively controlled".

2-loop Casimir stabilization

$$S^1/(z_2 \times z_2'): \quad V(R) \sim \frac{1}{R^4} + \frac{g^2}{R^5} \ln(\Lambda \cdot R)$$

\uparrow
 cutoff

Origin of log-enhancement

- At 1-loop, brane-localized kinetic terms (e.g. F^2) arise
- They are UV-divergent and modify KK-spectrum by terms $\sim g^2$
(\rightarrow Barbieri/Hall/Nomura & many others)
- Calculating the 1-loop Casimir energy with these corrections gives the above "dominant 2-loop effect" (indep. of the UV completion)
- We choose $\Lambda = \Lambda_{\max} = M \quad \Rightarrow \quad V(R) \sim \frac{1}{R^4} + \frac{\ln(MR)}{MR^5}$

SUSY-Breaking

- Of course, $V(R) \equiv 0$ if SUSY is unbroken.
- We assume Scherk-Schwarz SUSY-breaking with an $SU(2)_R$ -twist $\omega \ll 1$. ($m_{1/2} \sim \omega/R$)
- Equivalently, we can think of a no-scale model with radion superfield $T = R + \dots$. In the presence of $W = W_0$, we find $F_T \neq 0$ ($F_T \sim \omega \sim \frac{|W_0|}{M_{P,5}^3}$)

(cf. Marti/Pomarol; Kaplan/Weiner;

Chacko/Luty; Gersdorff/Quiros/Riotto, ...)

- Since both terms in $V(R) \sim \omega^2$, the SUSY-breaking scale is irrelevant for the position of the minimum

The Loop Calculation

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needs careful book-keeping, but can be done without new integrals (by thoughtful use of (N=2) SUSY, elementary group theory and \mathbb{Z}_2 -trf. properties).

One finds e.g. (for $S^1/(\mathbb{Z}_2 \times \mathbb{Z}'_2)$, with $G \rightarrow H$ at one boundary)

$$C^{(1)} \sim \omega^2 (24 d_G - 56 d_H) \\ + \text{hypers} + \text{gravity}$$

$$C^{(2)} \sim \omega^2 \frac{2}{\pi^3} C_G (7 d_G - 15 d_H)$$

$$+ \text{hypers} \quad (\text{no gravity since } M \ll M_{Pl,5})$$

(Note: Calculation can be partially checked against Cheng/Matchev/Schmeltz: "Radiative corrections to KK-masses")

5d SUSY-GUT models

- $S^1 / (Z_2 \times Z_2')$; $SU_5 \rightarrow SU_3 \times SU_2 \times U_1$ by Z_2'
- Higgs on SM-brane
- Matter from "split multiplets" in bulk or on SM-brane

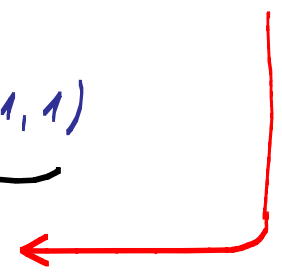
In detail: $\bar{5}$ -matter

$$\bar{5} = \underbrace{(\bar{3}, 1)}_u + \underbrace{(1, 2)}_s$$

← counts # of bulk fields of this type

10-matter

$$10 = \underbrace{(3, 2)}_t + \underbrace{(\bar{3}, 1) + (1, 1)}_r$$



$\Rightarrow r, s, t, u \in (0, 1, 2, 3)$

parameterize our "micro landscape" of $4^4 = 256$ models

• We find:

$$\frac{C^{(1)}}{C^{(2)}} = \pi^3 \frac{-160 - 16r - 8s + 96t + 48u}{\frac{2424}{5} - 108r - 36s + 36t + 12u}$$

• From this ratio,

the value of R_{\min} and hence α_{GUT} follow.

(We need $C^{(1)} < 0$, $C^{(2)} > 0$, and small ratio to get large MR.)

We find a minimum at $MR \approx 10$ for $\sim 1/3$ of the models

We find $\frac{1}{20} \approx \alpha_{GUT} \approx \frac{1}{30}$ for 12 models

Uplifting

$N=1$ supergravity perspective:

$$\Omega \sim - (T + \bar{T}) + \Delta\Omega_{\text{loop}} \quad ; \quad W = W_0$$

$$\Delta\Omega_{\text{loop}} \sim \frac{1}{(T + \bar{T})^2} + \frac{-g^2}{(T + \bar{T})^3} \ln(M(T + \bar{T}))$$

(cf. "Almost no-scale" proposal of Luty/Okada)

$$V(R_{\min}) \sim \frac{\omega^2}{R_{\min}^4} \text{ is calculable \& } < 0.$$

To "uplift", allow for a small warping ($\Lambda_5 \neq 0$):

$$\Lambda_5 = -6k^2 M_{P,5}^3 \quad ; \quad e^{-k(T + \bar{T})} = 1 - \text{"small"}$$

- Neglecting the loop-induced Kähler-correction for the moment, we have:

$$\Omega \sim e^{-k(T+\bar{T})} - 1 \quad ; \quad W = W_0 e^{-3kT}$$

(\rightarrow Luty/Sundrum)

(constant superpotential
at IR-brane)

- A Kähler-Weyl rescaling gives

$$\Omega \sim 1 - e^{k(T+\bar{T})} \quad ; \quad W = W_0$$

$$\sim -(T+\bar{T}) + \Delta\Omega_w$$



$$\Delta\Omega_w \sim (T+\bar{T})^2$$

- This has to be combined with our $\Delta\Omega_{loop}$

$$\Rightarrow \Delta\Omega_{tot} \sim -M_{P,5}^3 k (T+\bar{T})^2 + \frac{1}{(T+\bar{T})^2} + \frac{g^2}{(T+\bar{T})^3} \ln(M(T+\bar{T})) \quad 12$$



$$\delta V \sim \frac{|W_0|^2}{M_{P,5}^6} (\Delta\Omega)_{T\bar{T}}$$

$$\delta V \sim \underline{\underline{\text{const.}}} + \frac{1}{R^4} + \frac{g^2}{R^5} \ln(MR)$$

positive
constant
contribution $\sim \omega^2 M_{P,5}^3 k$

- Uplifting to $\Lambda_4 \approx 0$ implies

$$k(T+\bar{T}) \sim \frac{1}{(RM_{P,5})^3}$$

i.e. Warping is indeed small and our unwarped calculation is Ok.

Note:

We have checked explicitly that this is equivalent to "detuning" the IR-brane-tension à la Bagger/Belyaev

(cf. Bagger/Redi '03 ; Falkowski '05 ; Katz/Redi/Shadmi/Shirman '05 ; Gersdorff/Quiros/Riotto '03 ; Maru/Sakai/Uekusa '06)

In this approach, one simply modifies the (negative) IR-brane tension w.r.t. its RS-model-value.

Note:

If one insists on $\Lambda_5 = 0$ (e.g. because one doubts that $\Lambda_5 \neq 0$ can be realized in the heterotic setting), another option is a brane-localized F-term uplift

(cf. Gomez-Reino/Scrucça ; Lebedev/Nilles/Ratz ; Choi/Jeong ; Brümmer/A.H./Trapletti ; Dudas/Mambrini ; ...)

This is consistent with our calculation, but not very elegant...

Summary / Conclusions

- Two-loop Casimir stabilization occurs naturally in 5d orbifolds (given only F_T -dominance & suitable matter content).
- With a very mild discrete tuning, it can "explain" the smallness of $\alpha_{\text{GUT}} \sim 1/25$
- An elegant way of uplifting the stabilized AdS_4 -vacua is to appeal to a small warping ($\Lambda_5 \neq 0$)
- It would be interesting to find 2-loop-Casimir-stabil. in explicit heterotic models, in particular to allow for more moduli
(cf. e.g. 6d analysis of Buchmüller / Catena / Schmidt-Hoberg)